



# LuGRE

LUNAR GNSS RECEIVER EXPERIMENT

## The Lunar GNSS Receiver Experiment (LuGRE)

Joel J. K. Parker, NASA Goddard Space Flight Center  
on behalf of the LuGRE team

January 26, 2022



# The LuGRE Team



## NASA

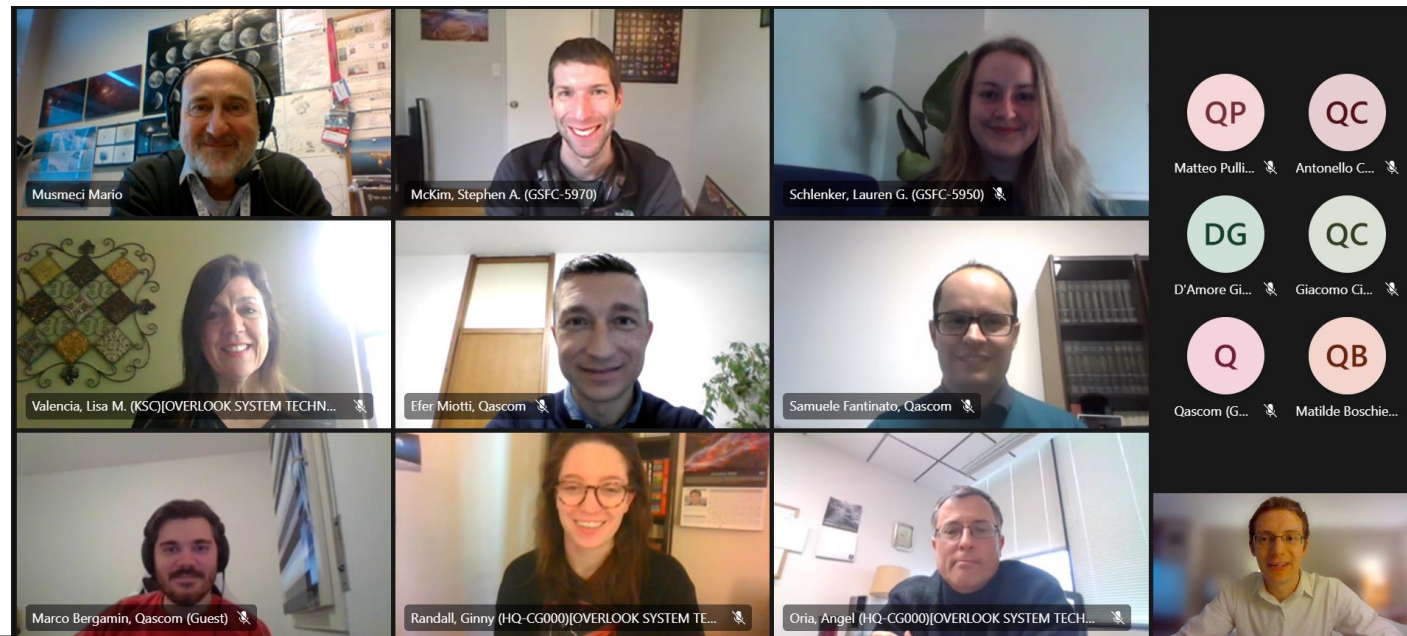
- **PI:** Joel Parker, NASA GSFC
- **Mission Manager:** Ben Anderson, NASA GSFC
- **Systems Engineer:** Steve McKim, NASA GSFC
- **Sponsor:** JJ Miller, NASA SCA
- **Advisors:** Lisa Valencia, Frank Bauer

## Italian Space Agency (ASI)

- **Co-PI:** Fabio Dovis
- **Project Manager:** Claudia Facchinetti
- **System Manager:** Mario Musmeci
- **Sponsor:** Alberto Tuozi
- **Technical Team:** Luigi Ansalone, Giuseppe D'Amore, Gabriele Impresario

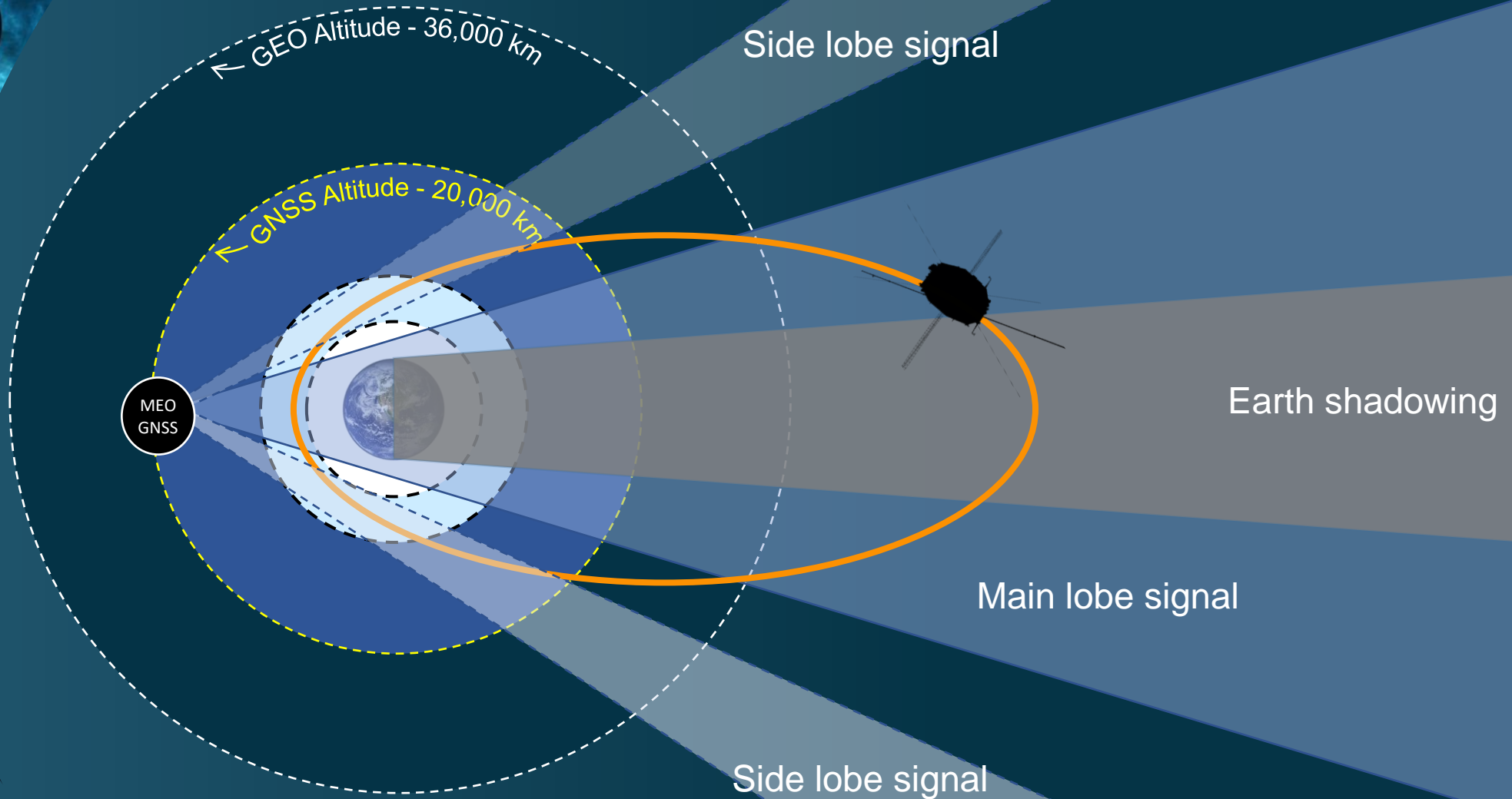
## Qascom, S.r.l

- **Project Manager:** Samuele Fantinato
- **Systems Engineering Manager:** Efer Miotti
- **Technical Director:** Oscar Pozzobon

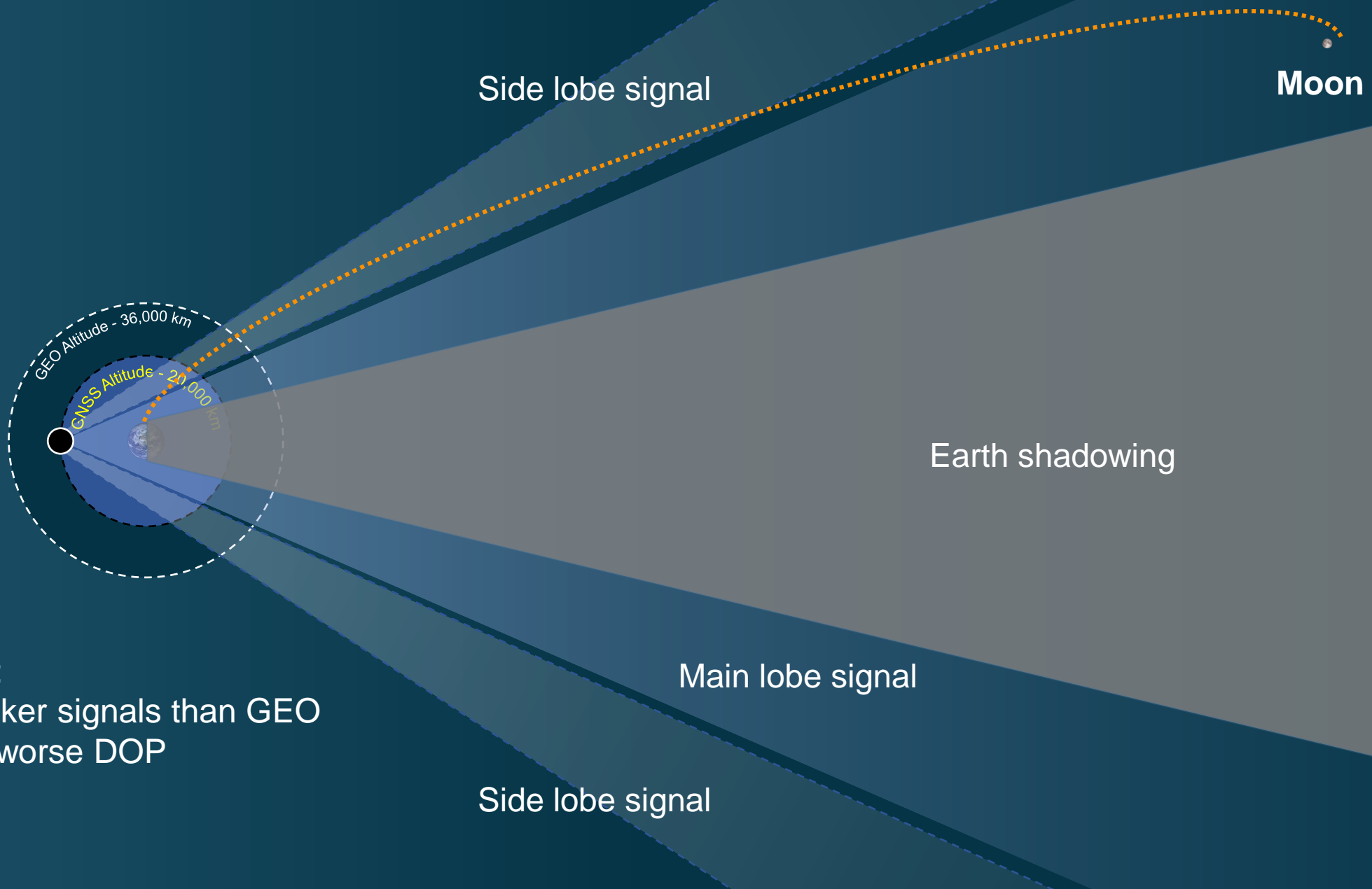




# Signal Reception in the Space Service Volume (SSV)



# Signal Reception **beyond** the Space Service Volume (SSV)



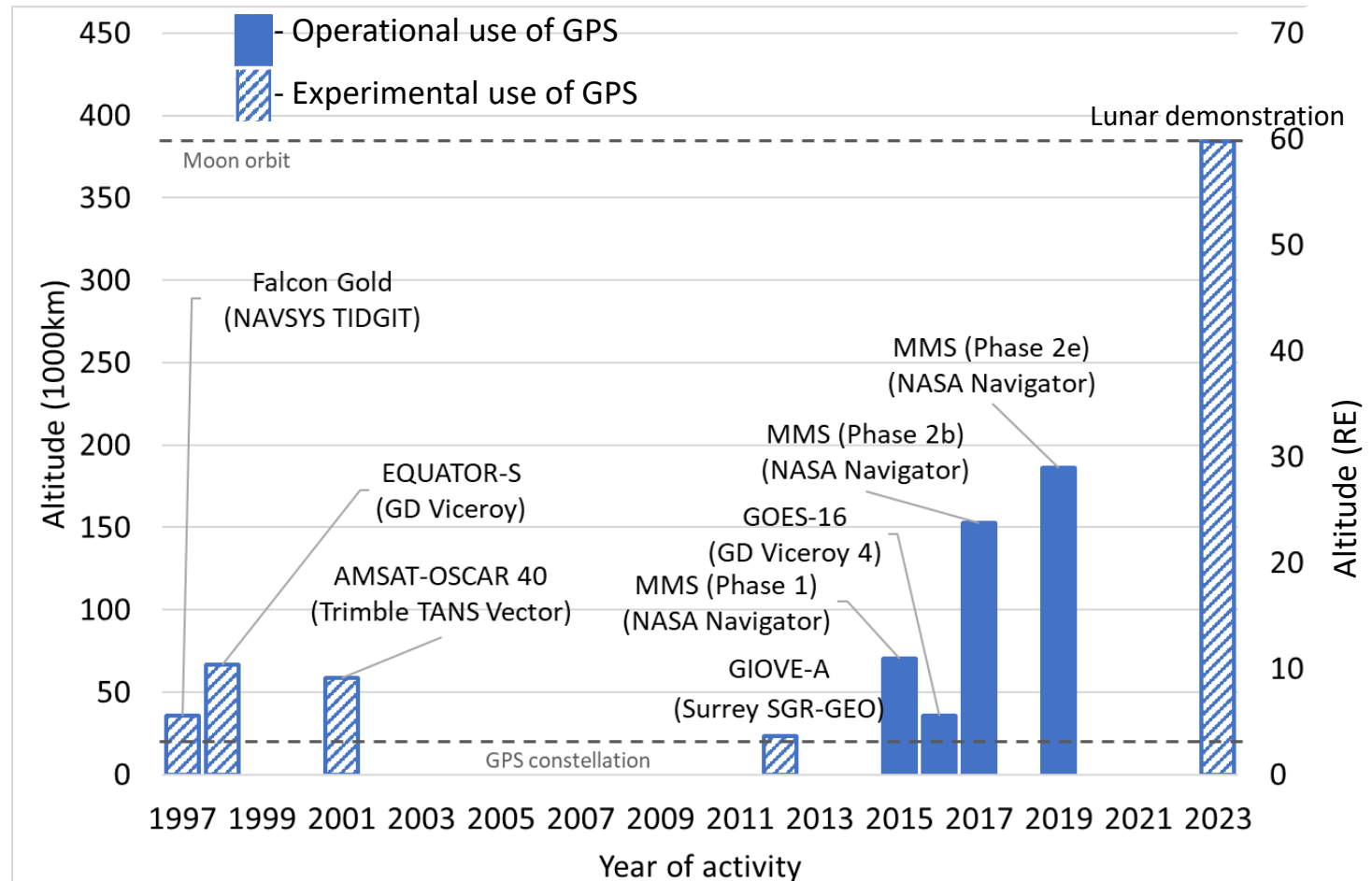
## Challenges:

- >30x weaker signals than GEO
- 10–100x worse DOP

# Development of High Altitude GNSS

## Transition from experimentation to operational use, and move into cislunar space:

- **1990s:** Early flight experiments—Equator-S, Falcon Gold
- **2000:** Reliable GPS at GEO w/ bent pipe architecture
- **2001:** AMSAT OSCAR-40 mapped GPS main and sidelobe signals
- **2015:** MMS employed GPS operationally at 76,000 km
- **2016–2017:** GOES-16/17 employed GPS operationally at GEO
- **2019:** MMS apogee raise to 50% lunar distance
- **2023: Lunar demonstration**



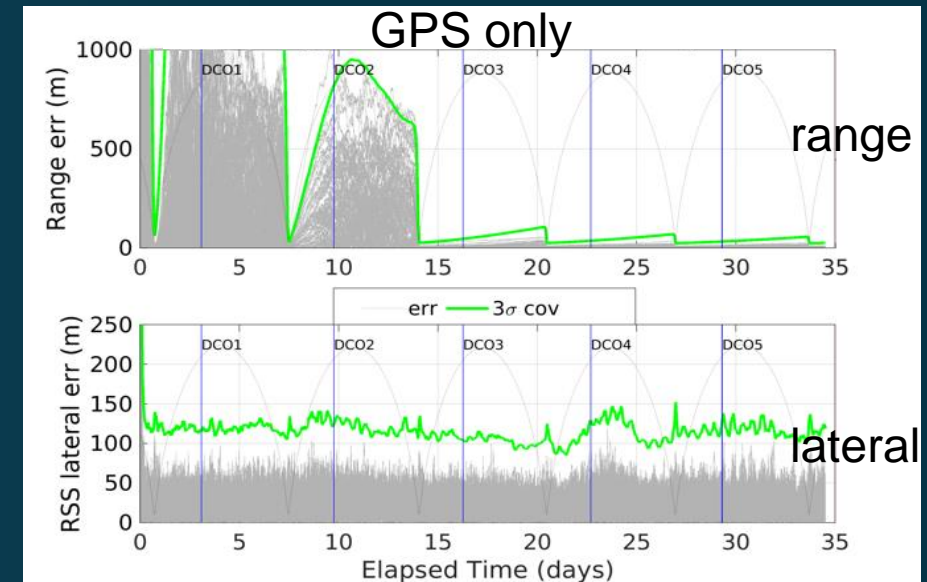
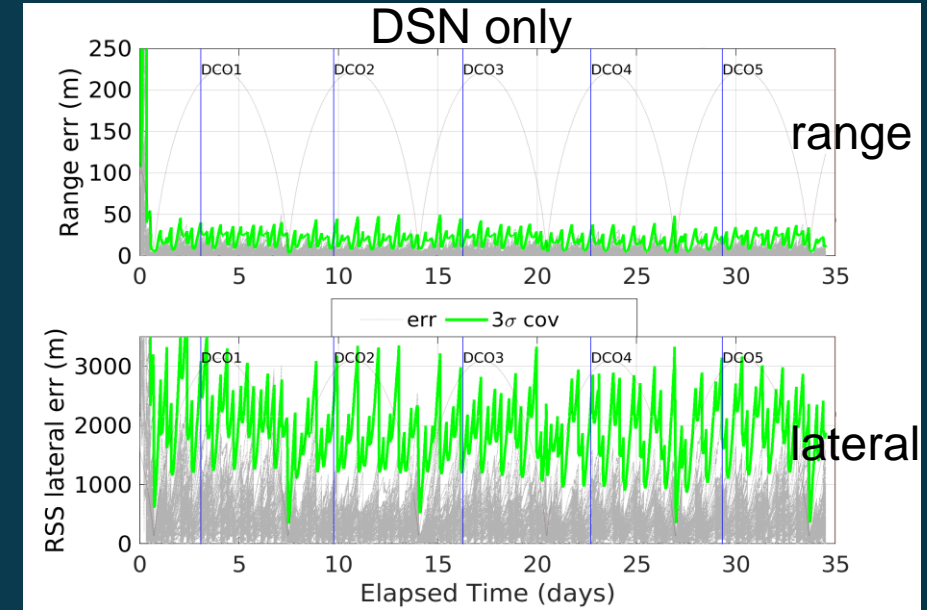
# Lunar Gateway Study – Sep 2020

## GPS Expected Performance

- Update to Feb 2019 preliminary study, both using MMS-calibrated models
- Position and velocity goals: 10 km and 10 cm/s, respectively
- Analyzed max OD error at the Data Cutoff (DCO) and at the final two perilunes and apolunes
- Observations:
  - GPS can provide greatly improved performance vs. DSN
  - GPS is real-time, on-board, without reliance on ground-based assets.

### Max steady-state errors, crewed assumptions

	Case	DCO	Apolune	Perilune	All
Position [m]	DSN	1469.7	1326.4	319.8	2353.6
	<b>GPS</b>	<b>60.4</b>	<b>84.5</b>	<b>73.0</b>	<b>118.7</b>
	DSN+GPS	57.7	81.7	107.0	117.4



# Lunar GNSS Phased Approach

## Initial demonstrations

- Demonstrate lunar reception
- Opportunistic flights/technology
- Return evidence of performance
- Return raw data, lessons learned

## First operational capability

- High-reliability unit
- High-accuracy clock
- Integrated into vehicle avionics
- Leverage demo data & lessons

## Commercialization

- Foster lunar rcvr commercial base
- Diversity of rcvr classes:
  - Flagship receivers
  - CubeSat/reduced SWaP
  - Integrated chipsets

## Broad Infusion

- GNSS is standard equipment
- Established, diverse rcvr base
- Part of diverse PNT solution

- GNSS-only baseline demonstrations

- Utilize initial lunar PNT services

- Leverage PNT signal compatibility:  
Earth-based GNSS + lunar PNT services

- Robust global PNT signal coverage from all sources



**LuGRE**

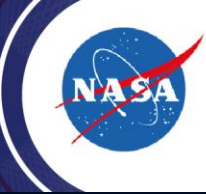
Relative use of signal sources

Terrestrial GNSS

Lunar PNT Services



# LuGRE Mission Overview



## Mission

- NASA HEOMD payload for CLPS “19D” flight
- Joint NASA/Italian Space Agency mission
- “Do No Harm” class
- Firefly Blue Ghost commercial lander
- Transit+surface observation campaign
- Expected surface duration: one lunar day (~12 Earth days)

## Payload objectives

1. Receive GNSS signals at the Moon. Return data and characterize the lunar GNSS signal environment.
2. Demonstrate navigation and time estimation using GNSS data collected at the Moon.
3. Utilize collected data to support development of GNSS receivers specific to lunar use.

## Measurements

- GPS+Galileo, L1/L5 (E1/E5)
- Onboard products: multi-GNSS point solutions, filter solutions
- Observables: pseudorange, carrier phase, raw baseband samples

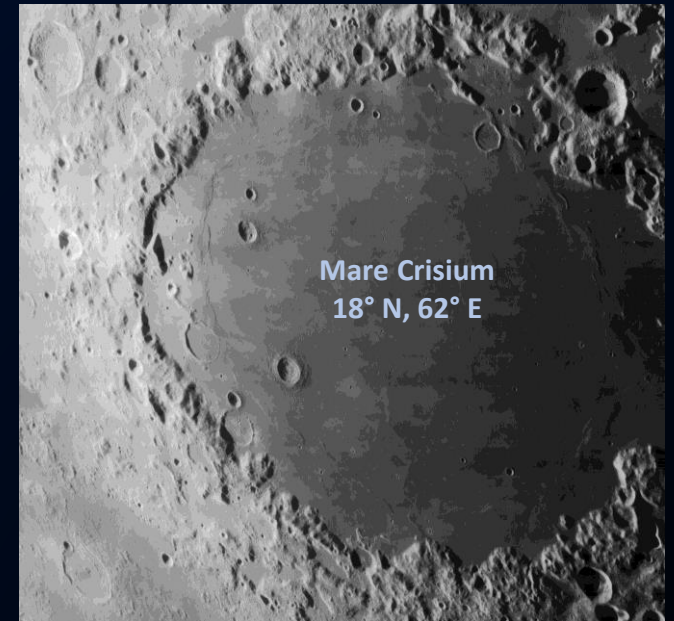
## Utilization

- Data + lessons learned for operational lunar receiver development
- Potential collaborative science: heliophysics, lunar geodesy
- Lunar human and robotic real-time onboard PNT

NASA

CLPS

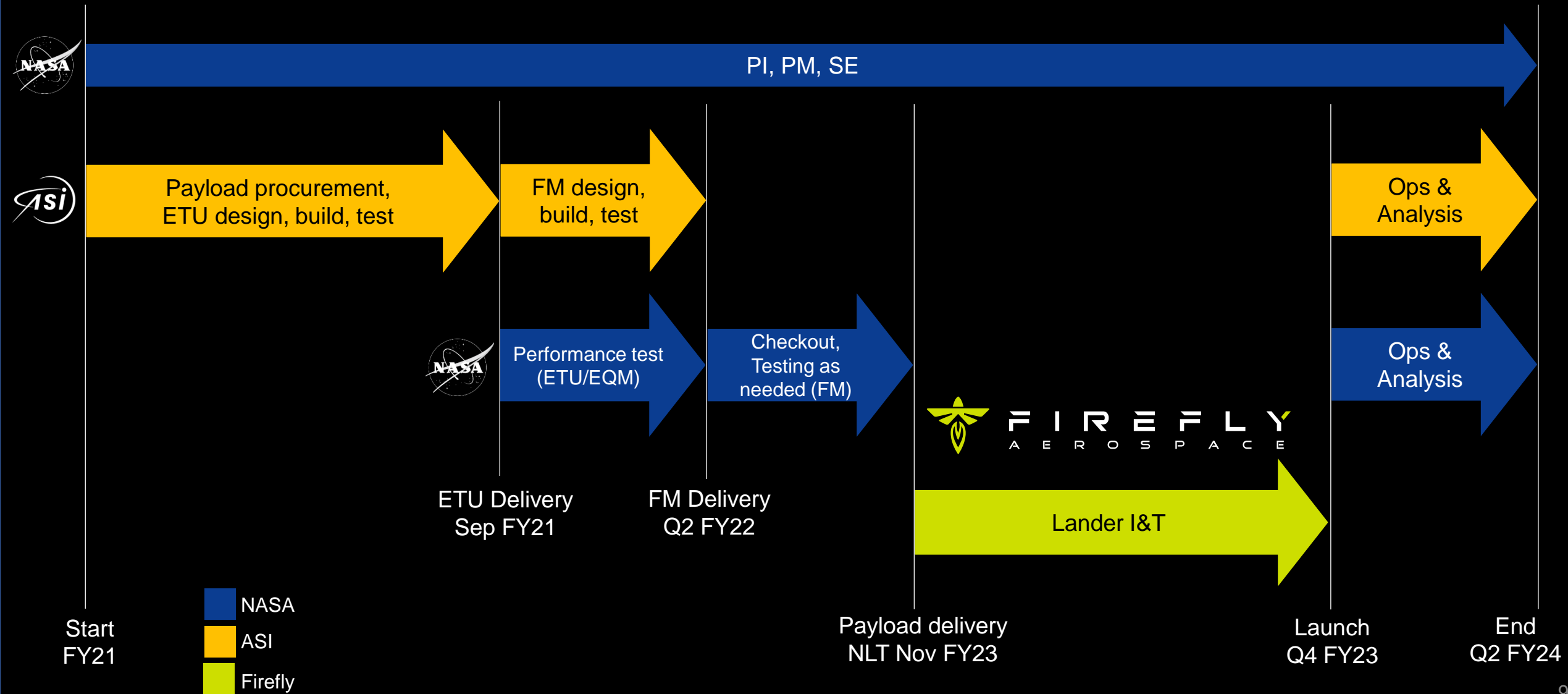
Commercial Lunar Payload Services



↑ We are here



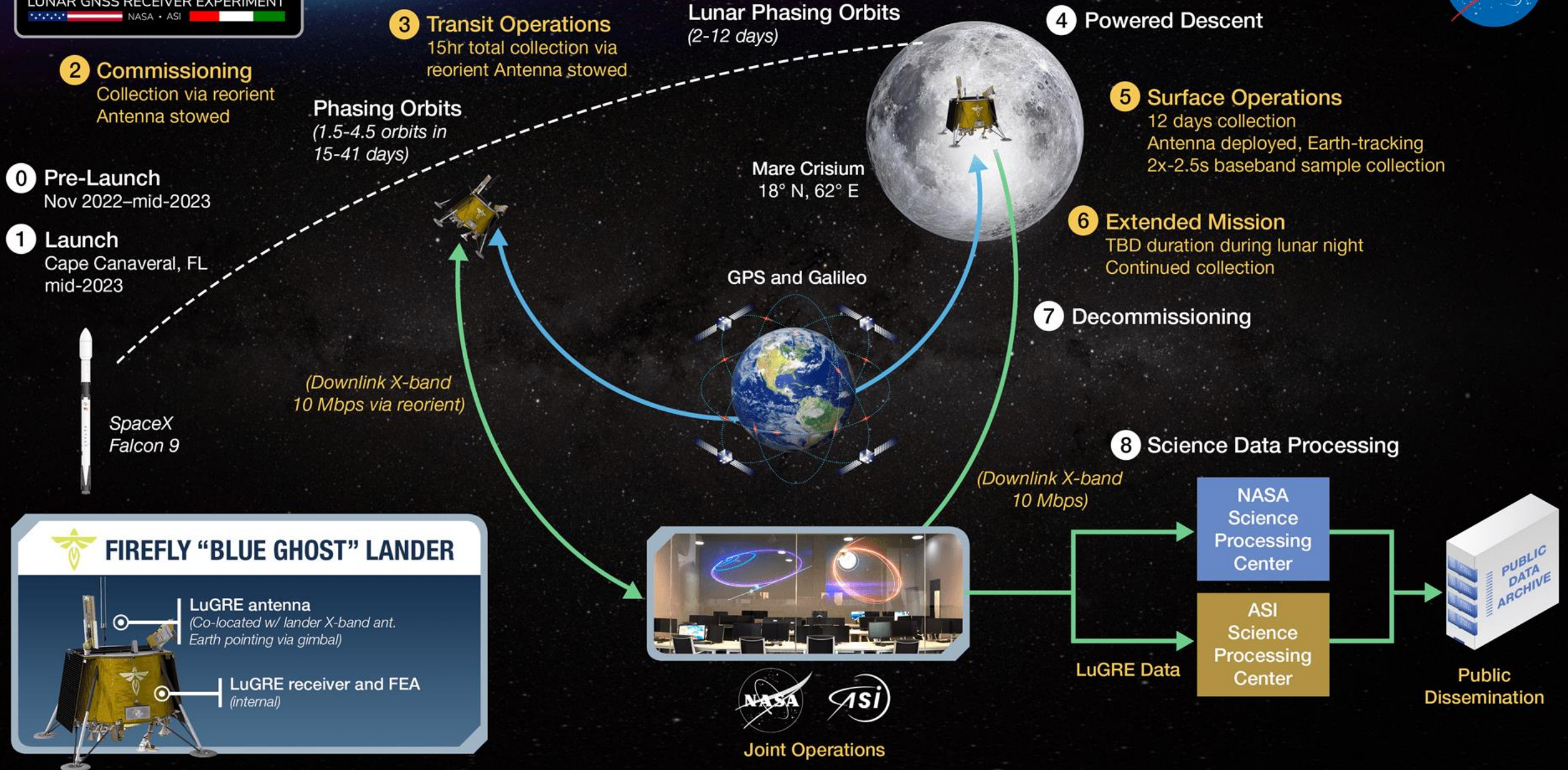
# LuGRE Roles



# LuGRE

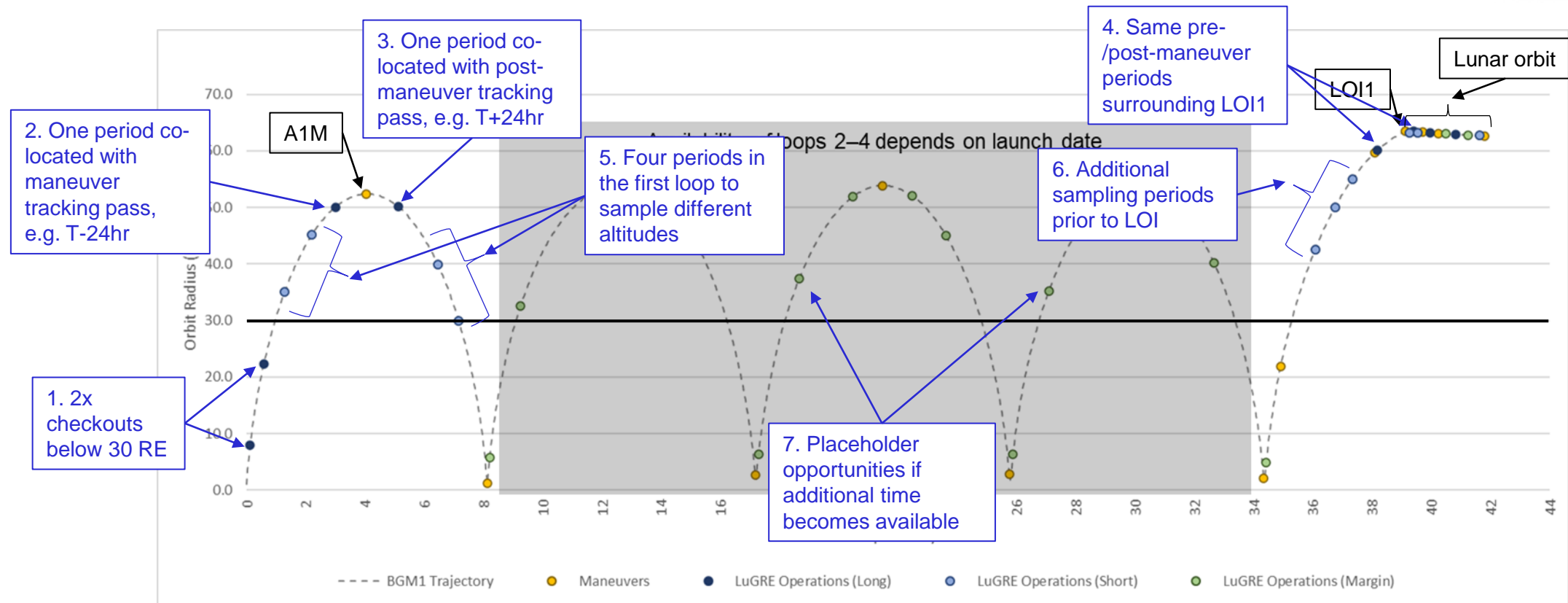
LUNAR GNSS RECEIVER EXPERIMENT

NASA • ASI





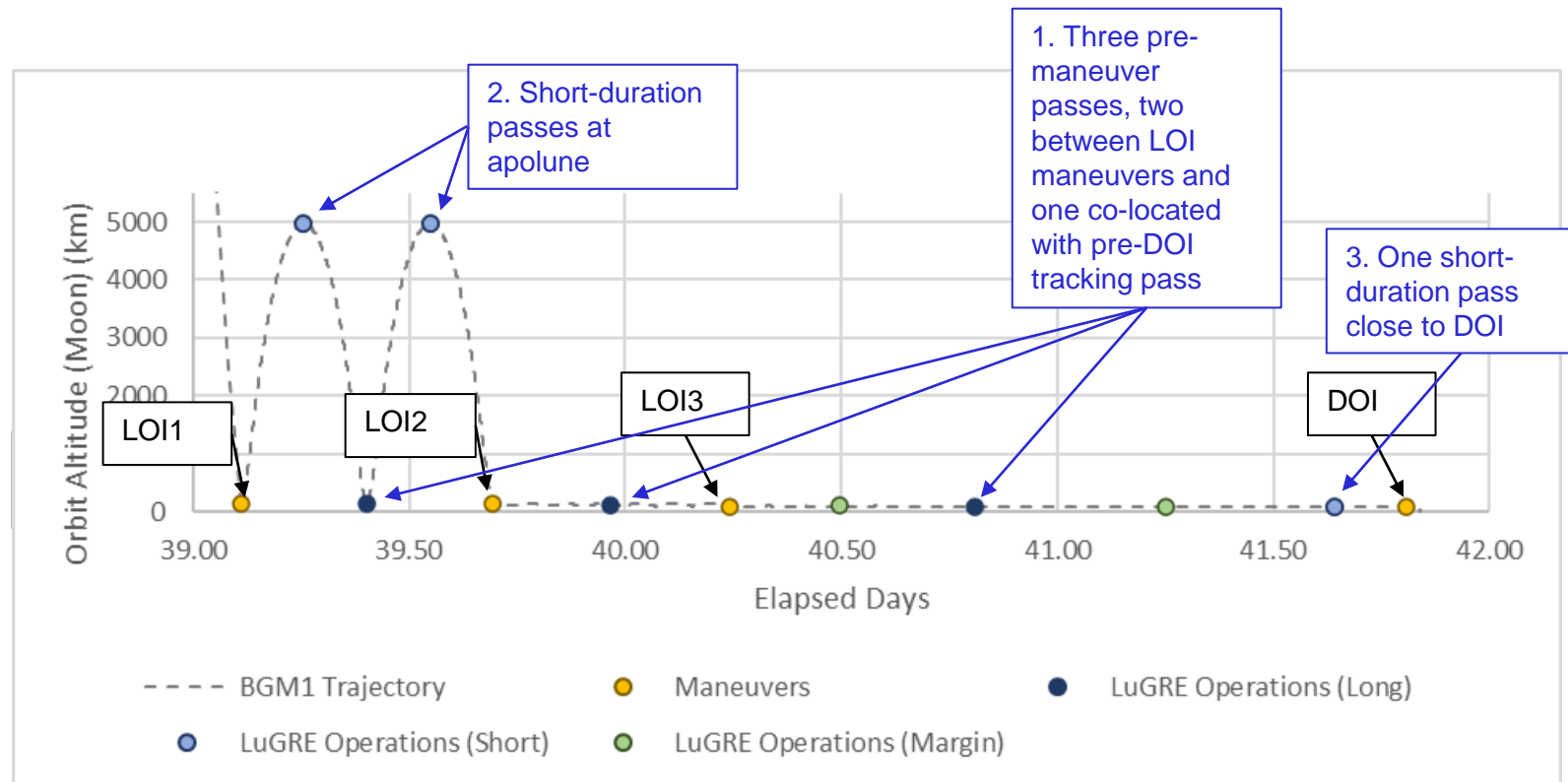
# Transit Operations Plan (Earth→Moon Transit)



- Initial “Full” operations plan, showing all features
- Total overall operations duration in transit is limited to 15hr as baseline due to limited ACS propellant available for Earth pointing
- Baseline (blue) operations have long (60 min) and short (42 min) durations; 18 total. Focus is altitude sampling and pre-/post- maneuver nav demo.
- Opportunistic (“margin”) operations (green) are pre-planned and can be activated if additional time becomes available



# Transit Operations Plan (Lunar)



- Lunar orbit portion: 5000km x 100km → 250km x 100km → 100km circular
- Total duration: 3 days
- Data collection focuses on key areas: demonstrate nav for maneuvers; leading into DOI; at perilune/apolune

# LuGRE Payload

## Payload characteristics

### 1. High-altitude GNSS receiver

- Qascom receiver, GPS+Galileo L1/E1 + L5/E5
- Based on QN400 flight heritage
- Cold redundant configuration
- Mass: 1.24 kg
- Power: 14 W

### 2. Low-noise amplifier (LNA)

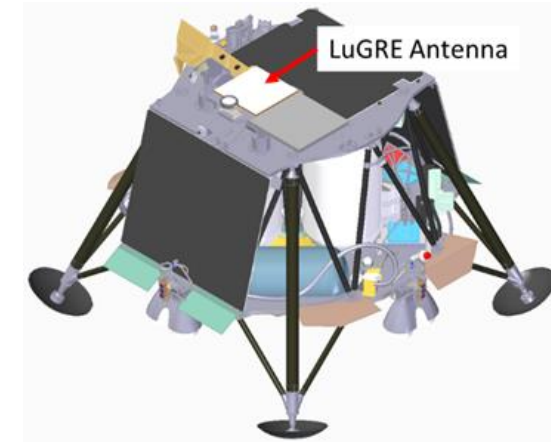
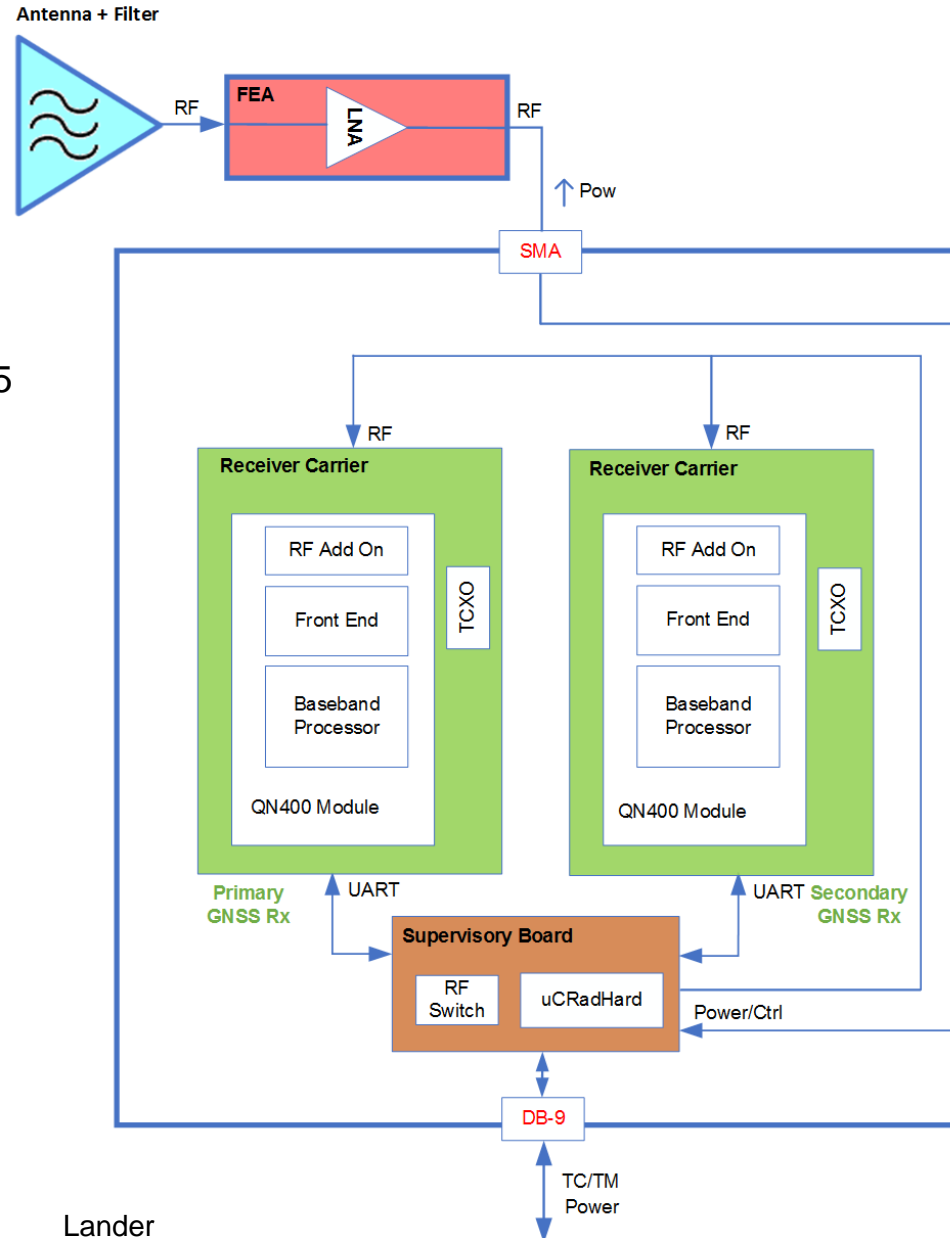
- Mass: 0.85 kg
- Power: 0.7 W

### 3. High gain L-band antenna + filter

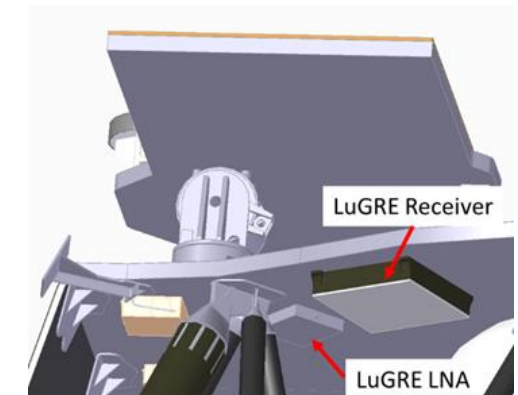
- Requires Earth pointing for GNSS reception
- 14 dBi peak gain, 10deg FOV
- Mass: 2.2 kg
- Power: 0 W (passive)

## Total resource allocations

- Mass: 4.64 kg
- Power: 14 W



Lander Top Deck, External View

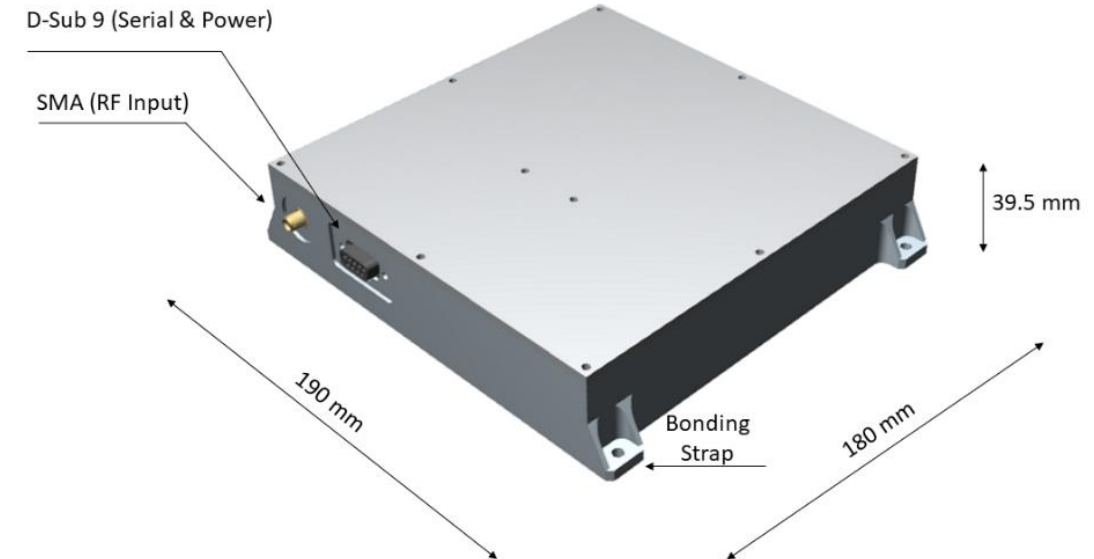


Lander Top Deck, Internal View

# Receiver



Parameter	Value
<b>Mass</b>	1.24 kg
<b>Power</b>	14 W (operating)
<b>Envelope</b>	19.0 x 18.0 x 3.95 cm
<b>Operating Temperature Range</b>	-35°C to 50°C
<b>Signal Reception</b>	GPS L1 C/A and L5 Galileo Open Service E1 and E5a
<b>Weak Signal Acquisition and Tracking Threshold</b>	< 23 dB-Hz
<b>Capabilities</b>	<ul style="list-style-type: none"> <li>Lunar-capable extended Kalman filter</li> <li>Capture of raw IQ samples</li> <li>Navigation ephemeris and aiding data upload via telecommand</li> </ul>
<b>Data Product Output</b>	<ul style="list-style-type: none"> <li>Least-squares point solutions</li> <li>Extended Kalman filter solutions with covariance</li> <li>Pseudorange observations</li> <li>Doppler</li> <li>Carrier phase observations</li> <li>Tracking status &amp; C/N0</li> <li>Raw IQ samples</li> </ul>



- Developed by Qascom, S.r.l. for ASI
- Based on Qascom QN400 flight heritage
  - Suborbital sounding rockets
  - Ohio University Bobcat-1 LEO CubeSat
- SDR architecture
- Dual-receiver cold-redundant configuration to mitigate single-event radiation effects



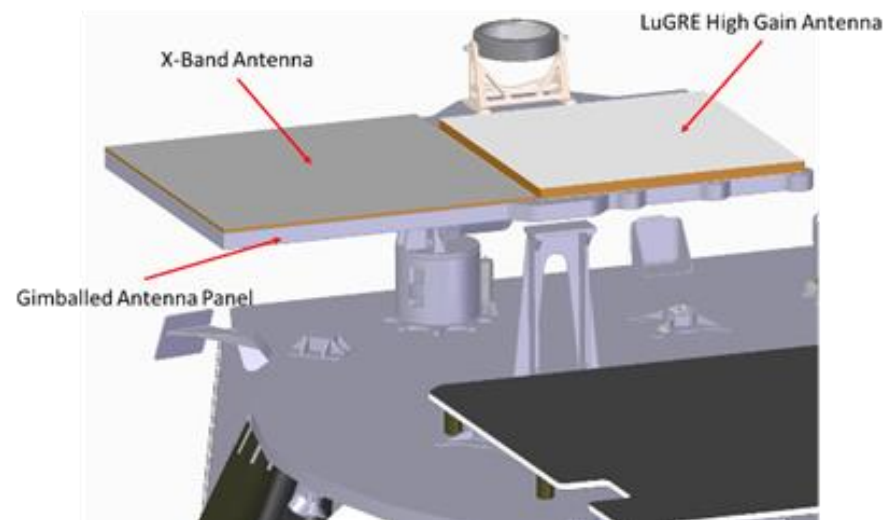
# HGA/LNA



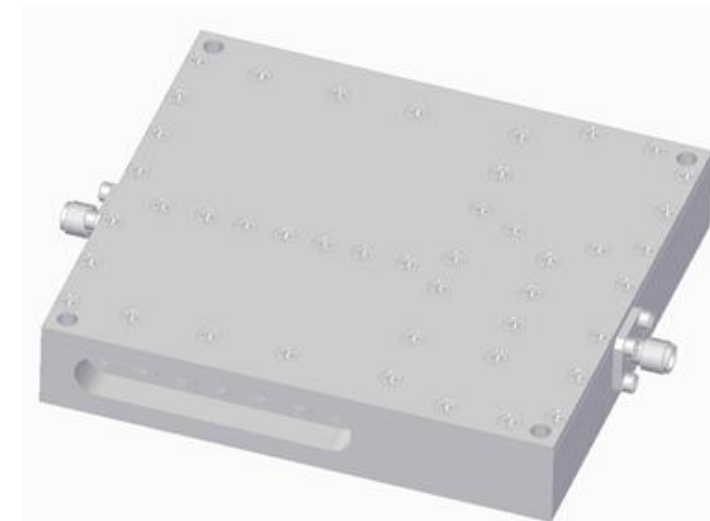
Parameter	Value
<b>Mass</b>	2.2 kg
<b>Power</b>	passive
<b>Envelope</b>	43.0 x 43.0 x 2.00 cm
<b>Operating Temperature Range</b>	-145°C to 125°C
<b>Antenna Type</b>	Passive Planar Antenna Array
<b>Polarization</b>	RHCP
<b>Gain</b>	≥ 14 dBi
<b>Working Band 1</b>	1575.42 +/- 12.276 MHz (L1/E1)
<b>Working Band 2</b>	1176.45 +/- 10.230 MHz (L5/E5a)
<b>Connector</b>	1x SMA

Parameter	Value
<b>Mass</b>	0.85 kg
<b>Power</b>	0.7W
<b>Envelope</b>	9.3 x 10.2 x 1.8 cm
<b>Operating Temperature Range</b>	-35°C to 50°C
<b>1<sup>st</sup> Band:</b>	1575.42 +/- 12.276 MHz (L1/E1)
<b>2<sup>nd</sup> Band:</b>	1176.45 +/- 10.230 MHz (L5/E5a)
<b>Noise Figure</b>	≤ 3 dB
<b>Connector</b>	2x SMA

HGA



LNA



# LuGRE Outcomes



Characterize the GNSS  
signal environment

- GPS+Galileo, L1+L5, E1+E5a
- Signal availability
- DOP
- $C/N_0$
- Observables
  - Pseudorange
  - Carrier phase
  - Doppler
- Raw baseband I/Q samples
- Transmit antenna patterns
- Multipath, surface environment

Characterize navigation  
performance

- Point solutions
- Onboard Kalman filter states
- Time to first position fix
- Formal errors, convergence
- Comparison to independent sources (lander, LRR)
- Application of GGTO

Share collected data

- GNSS receiver developers
- LuGRE science partners
- NASA missions (Artemis, Gateway, science)
- Commercial landers
- International space agencies
- GNSS community
- Science community
- Public

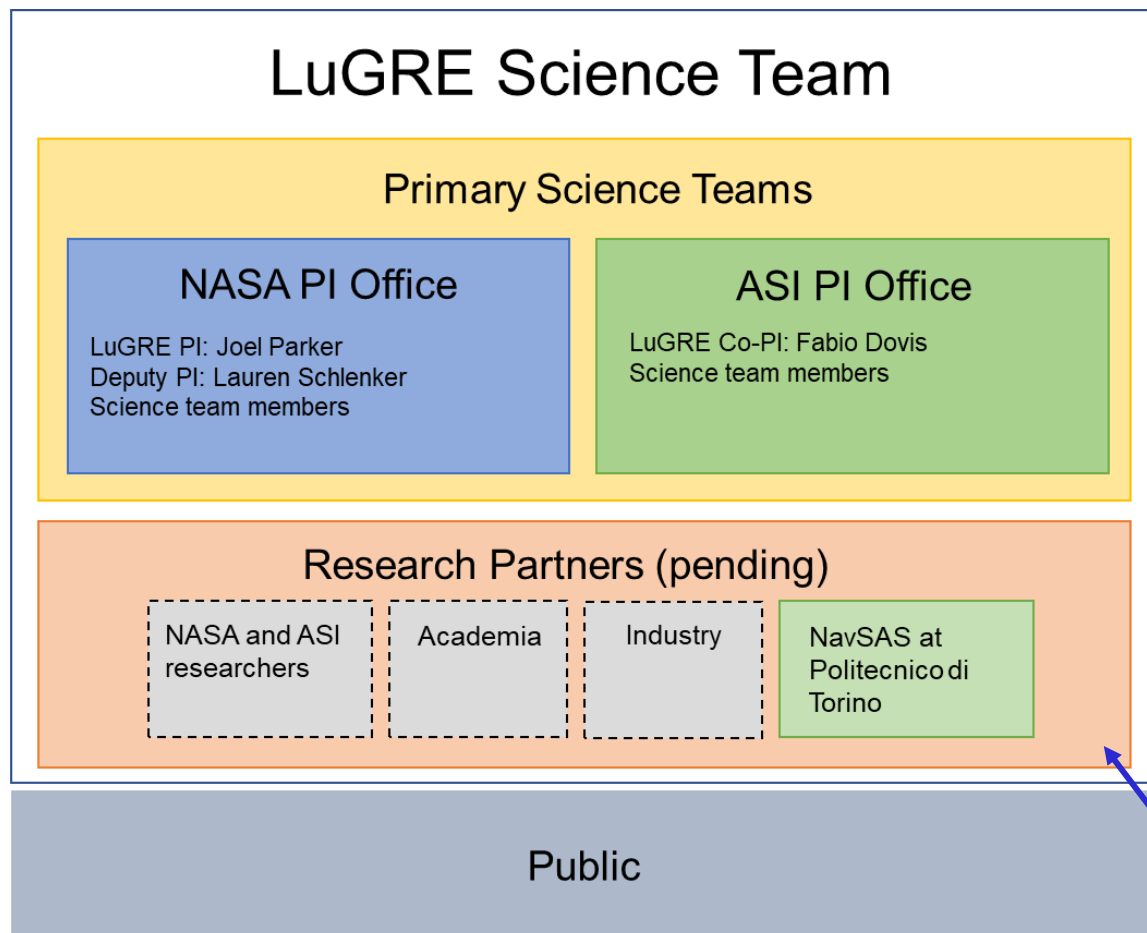
Facilitate adoption of  
capability

- Raw data availability
- LuGRE team reports + papers
- Calibration of lunar GNSS simulation models
- Application to future mission navigation studies
- Lessons learned to GNSS hardware and software developers

# Mission Science



## LuGRE Science Team structure:



## Driving investigations:

### Objective 1

- Measure the signal strength throughout the mission and empirically evaluate link budget model.
- Determine signal availability throughout the mission.
- Measure Doppler-shift and Doppler-rate profiles throughout the mission.
- Measure pseudorange from visible satellites during all planned operations periods.

### Objective 2

- Calculate and characterize least-squares multi-GNSS point solutions throughout the mission where sufficient signals are available.
- Calculate and characterize Kalman filter based navigation solutions onboard throughout the mission.
- Compare onboard navigation solutions to external sources (e.g., ground-based measurement processing, planned trajectory, Blue Ghost navigation solution).
- Characterize position, velocity, and time uncertainty and convergence properties throughout mission.

### Objective 3

- Process GNSS observables (e.g., Doppler, pseudorange) with ground-based tools to predict achievable onboard navigation performance.
- Calibrate ground models with LuGRE data and utilize to predict achievable navigation performance for future missions.

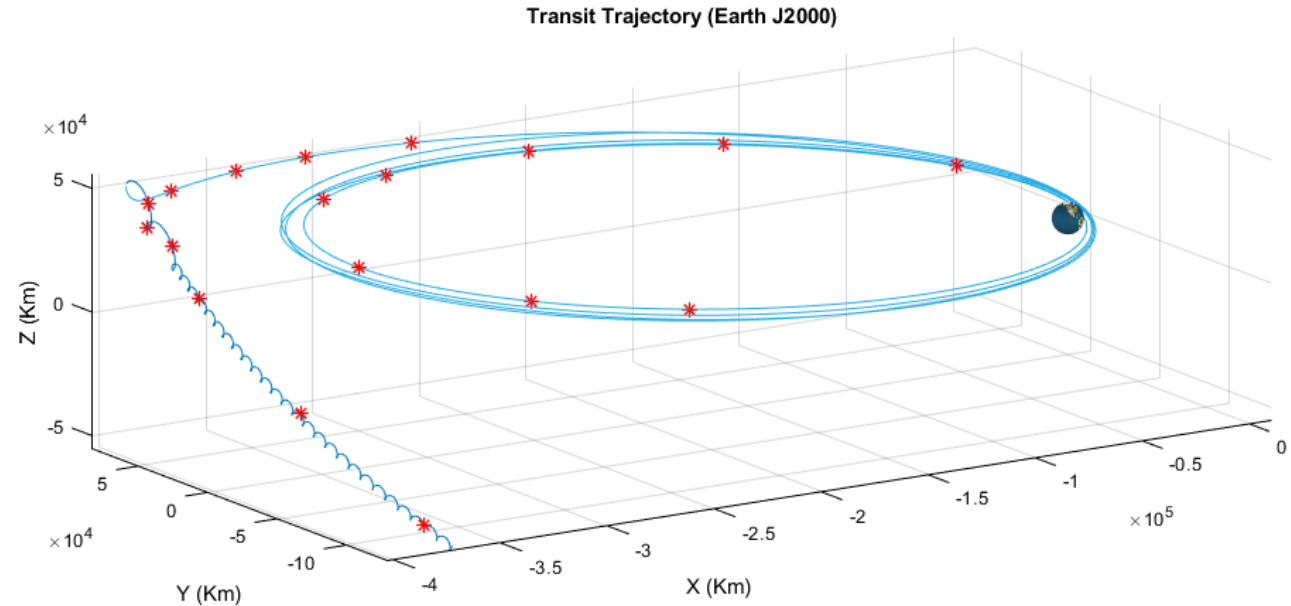
Opportunities for external research partners to get involved in science definition.



# Mission Performance Analysis

- Initial analysis of expected mission performance
- Focus: visibility (transit + lunar surface), real-time filter navigation performance
  - Baseline 18x operations periods during transit, continuous surface operations
  - L1 signal only, L5 simulation is future work
  - Receiver threshold 23 dB-Hz
- Tools/data:
  - NASA Goddard Enhanced Onboard Navigation System (GEONS)
  - Accessed via GEONS Ground MATLAB Simulation (GGMS)
  - High-fidelity GPS model calibrated with MMS flight data, GPS ACE transmit antenna patterns and available patterns from ground data
  - Publicly-available Galileo orbits, ESA 2019 antenna pattern
  - Firefly trajectory, Haigh-Farr simulated HGA pattern
- Link budget model:

$$C/N_0 = P_T + G_T(\phi, \theta) - 20 \log \left( \frac{4\pi d}{\lambda_{L1}} \right) + G_R(\phi, \theta) - L_{pol} - 10 \log(kT_{sys}) - R_{loss}$$

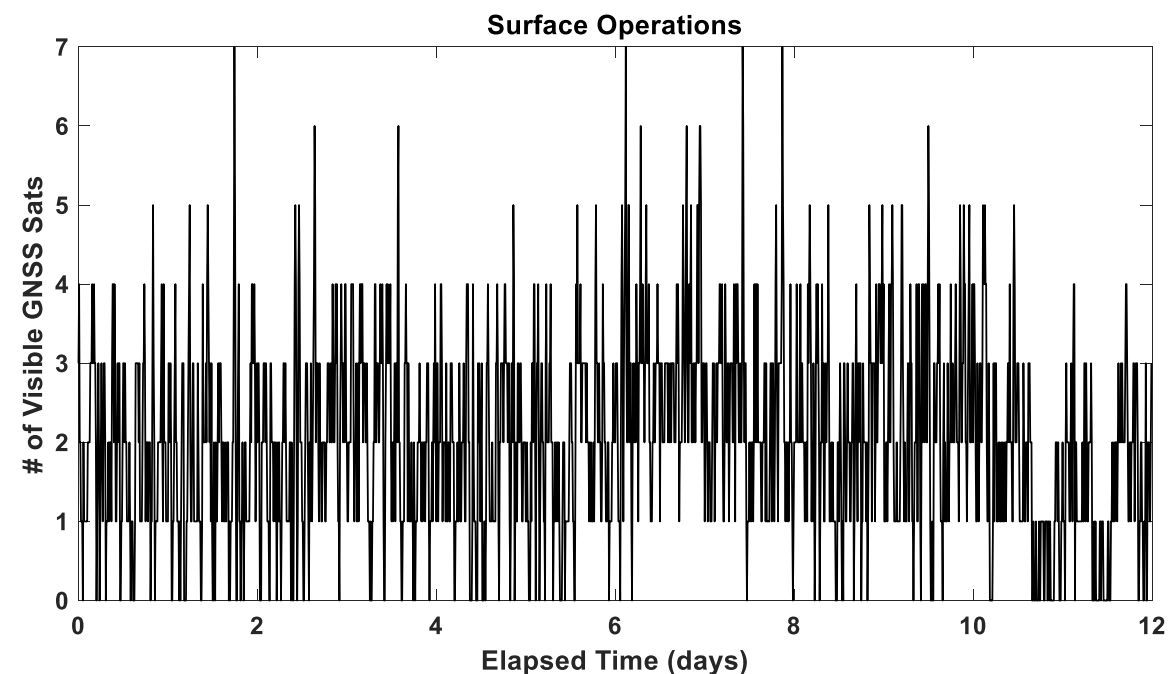
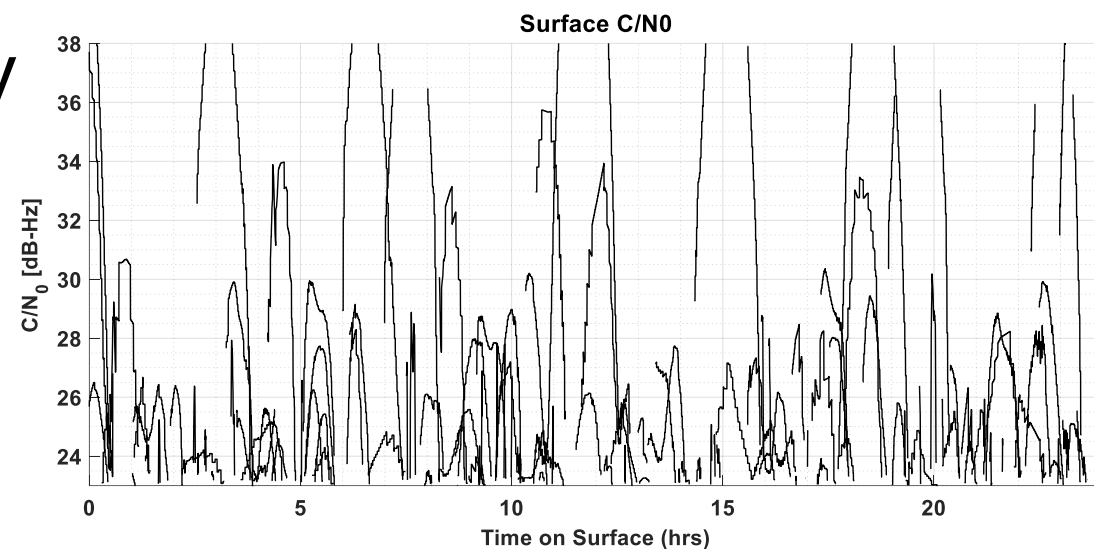
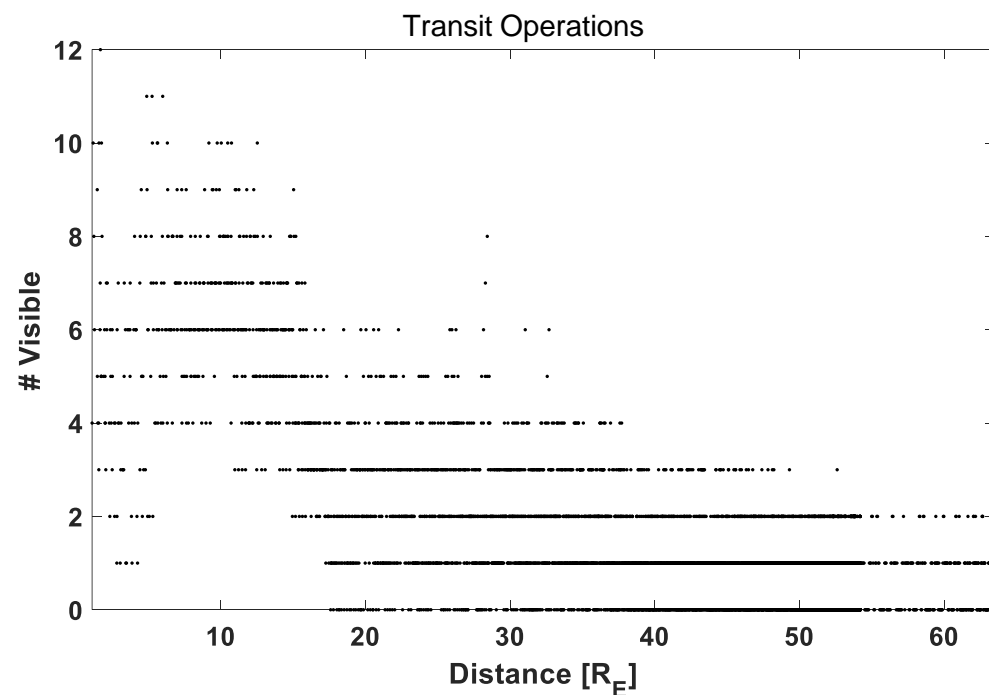


Link budget parameters

Parameter	Value
Receiver Implementation Losses $R_{loss}$	0.9 dB
System Temperature $T_{sys}$	295 K
Polarization Losses $L_{pol}$	3 dB
$P_T$ (GPS Block IIR)	17.3 dBW
$P_T$ (GPS Block IIR-M)	18.8 dBW
$P_T$ (GPS Block IIF)	16.2 dBW
$P_T$ (GPS Block III)	18.8 dBW
$P_T + G_T(\text{peak})$ (Galileo)	11 dBW
$G_R(\text{peak})$	16 dBW

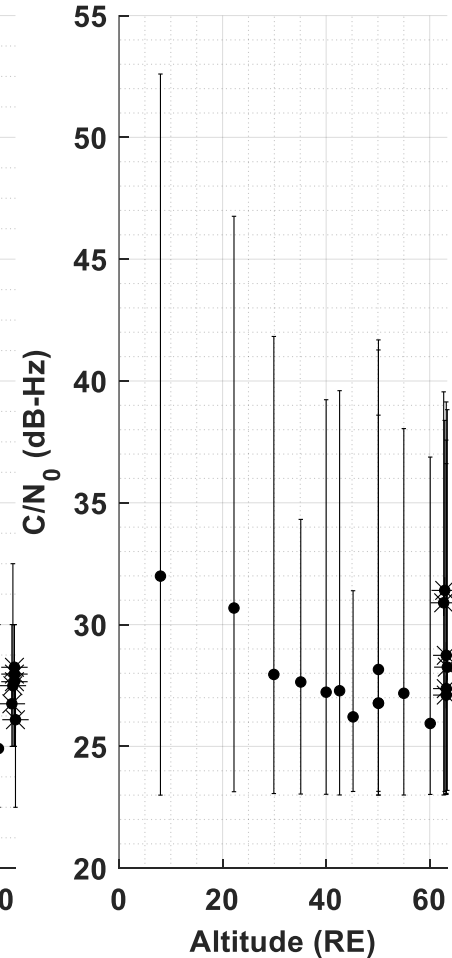
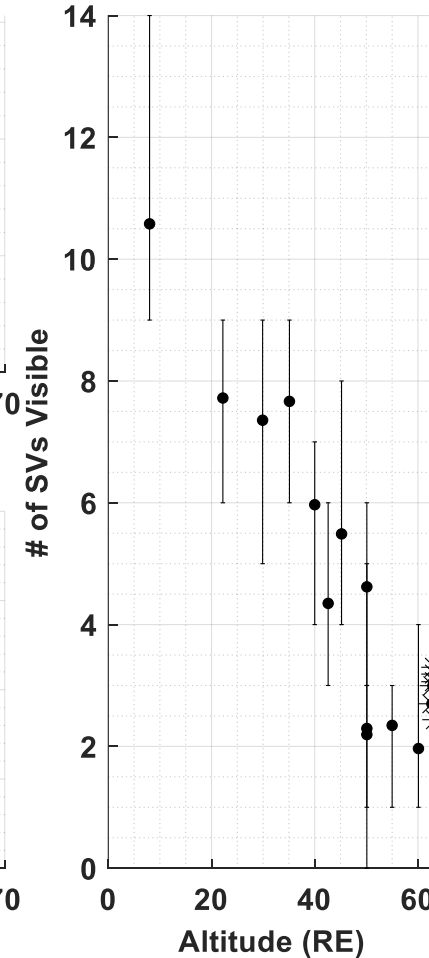
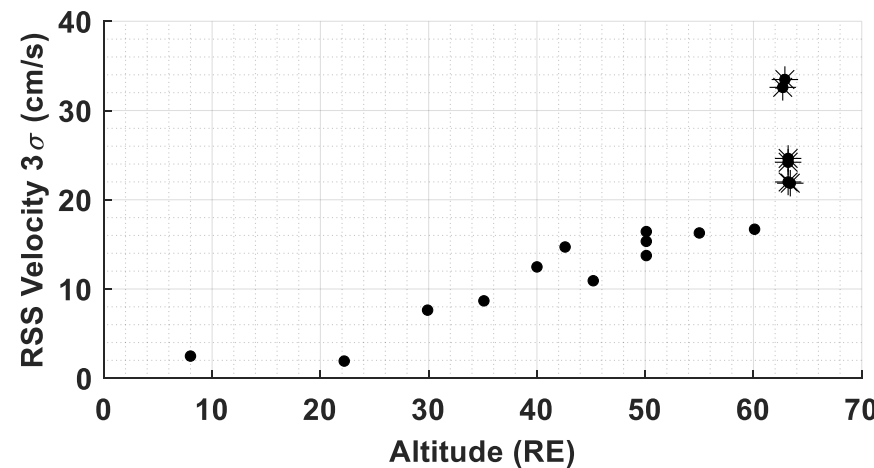
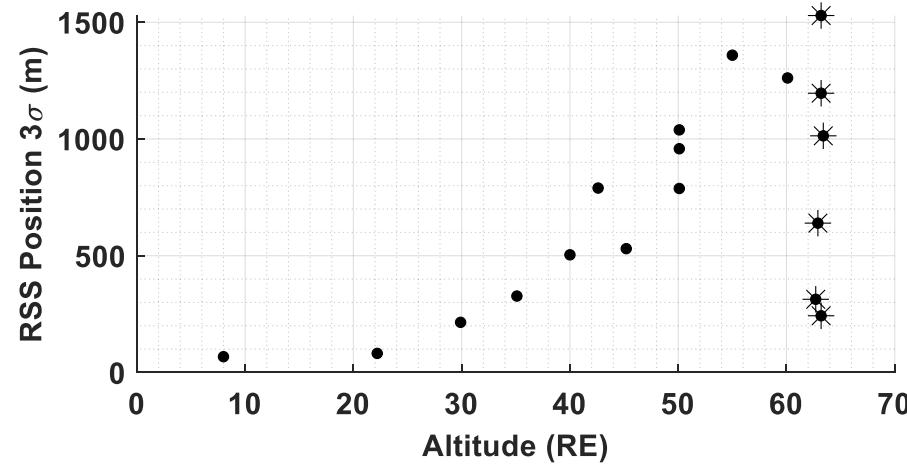
# Mission Performance: C/N0, Visibility

- Signal visibility extends to 60+ RE and is available on lunar surface.
- Average visibility at Moon is 2 SVs
- 4+ SV visibility achieved often in transit and at Moon (>10% of time)
  - Indicates occasional point solution availability
- Visibility is nearly all GPS, likely due to conservative Galileo assumptions
- C/N0 peaks >30 dB-Hz for main lobes, ~26 dB-Hz for side lobes



# Mission Performance: Transit Navigation

- NASA GEONS transit operations navigation performance simulation
- RSS position/velocity  $3\sigma$  covariance at end of operational arc (42 min or 60 min)
- Results generally a function of number of satellites visible as altitude increases.
- Final 6 periods in are in lunar orbit. Highly variable accuracy indicates sensitivity of solution to visibility and geometry.





# Mission Performance: Sample Capture

- Analysis by ASI with support by Politecnico di Torino
- Purpose: Assess feasibility of GNSS signal acquisition from raw IQ samples collected over a limited time window (ms)
- Result: Minimum coherent integration time needed to successfully acquire the signal, for different C/N0 and limiting the search space to Doppler bins
- $N_d$  = number of Doppler bins in search space

C/N0 (dB-Hz)	(ms) (Full SS)	(ms) ( $N_d = 5$ )	(ms) ( $N_d = 3$ )
36.4	6	4	4
32.2	8	6	6
27.2	55	45	45
24.0	120	90	90
21.9	105	100	100
20.2	175	160	160
18.6	415	370	230

# Conclusions

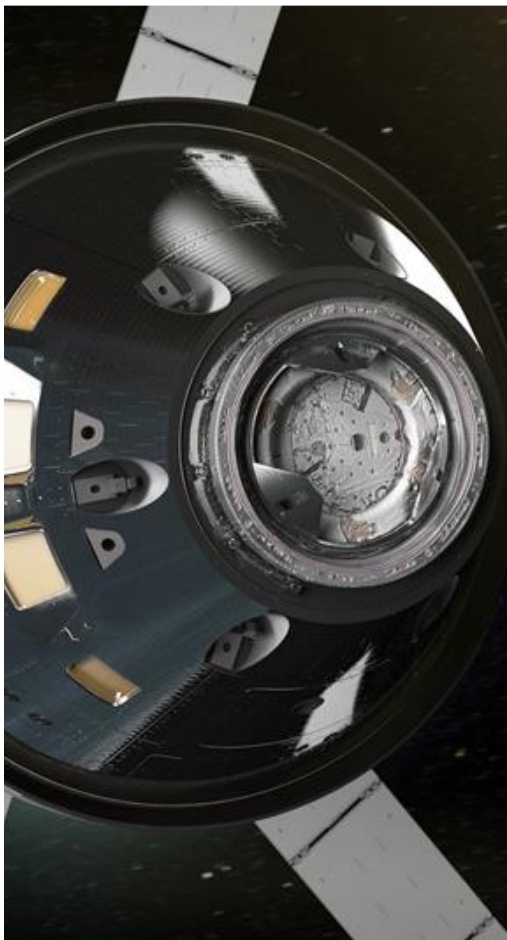
- LuGRE is a joint NASA/ASI project that will demonstrate GNSS-based PNT in transit to the Moon and on the lunar surface.
- LuGRE will fly a Qascom-designed weak-signal multi-GNSS receiver and high-gain antenna to receive GPS and Galileo signals in L1 and L5 bands.
- Initial performance simulations indicate a high degree of confidence in meeting the LuGRE science objectives to characterize the signal environment, perform navigation at the moon, and apply the results to future developments.
- **LuGRE will create an initial public dataset with the intention to jump-start development of GNSS-based navigation at the moon.**



# Backup



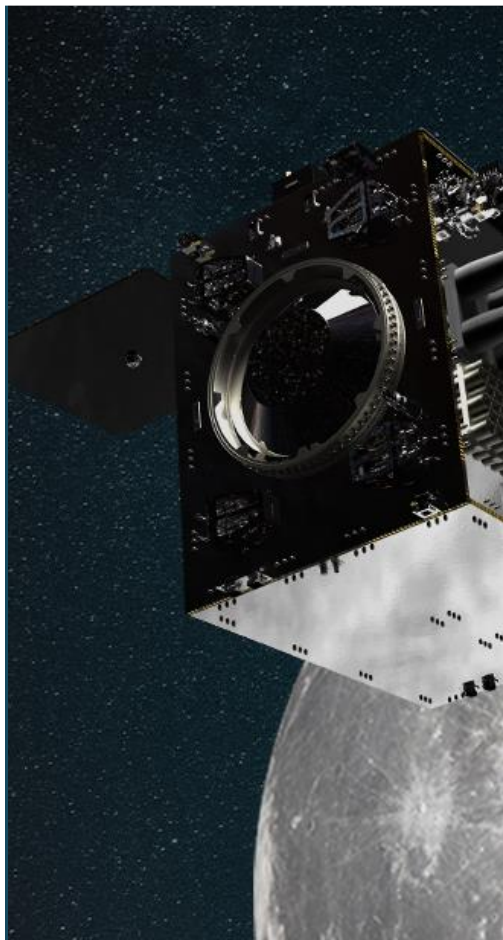
# What's Next



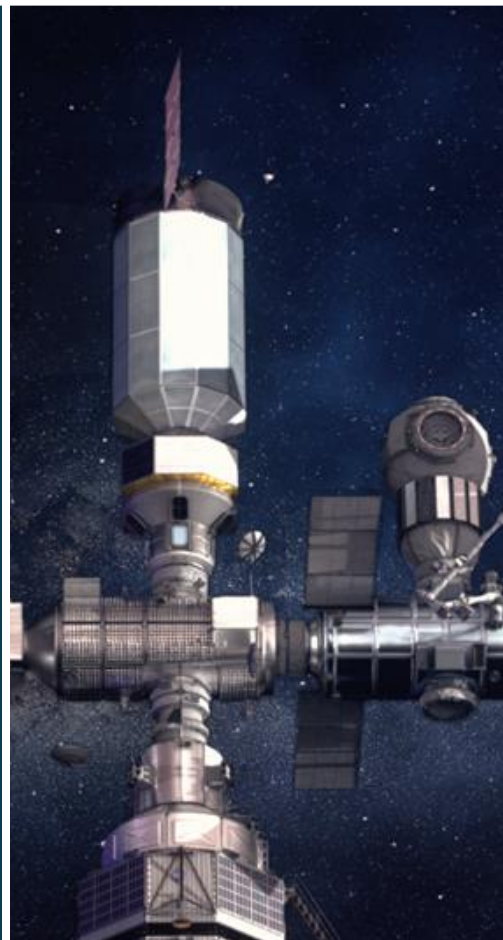
Artemis-1  
(LEO receiver)  
**2022**



LuGRE  
**2023**



Lunar Pathfinder  
(ESA)  
**2024**



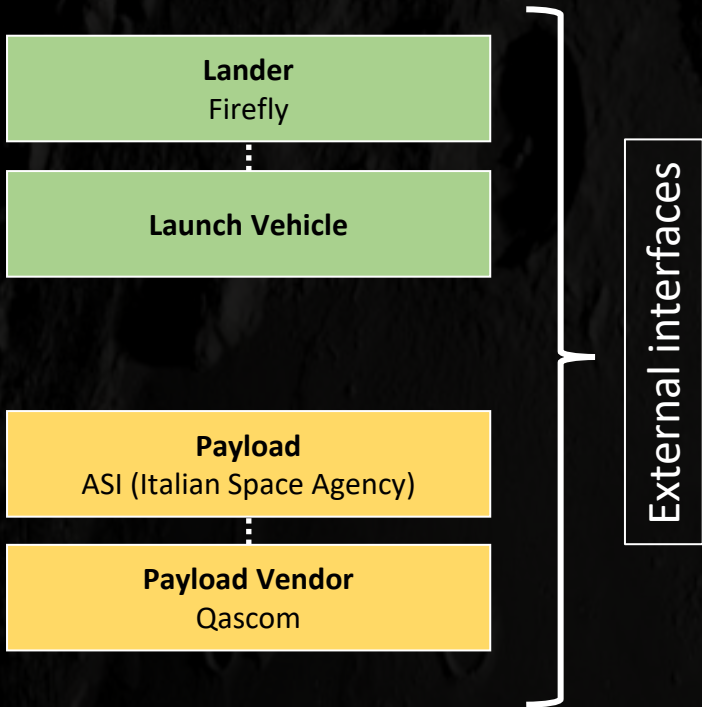
Gateway Payloads



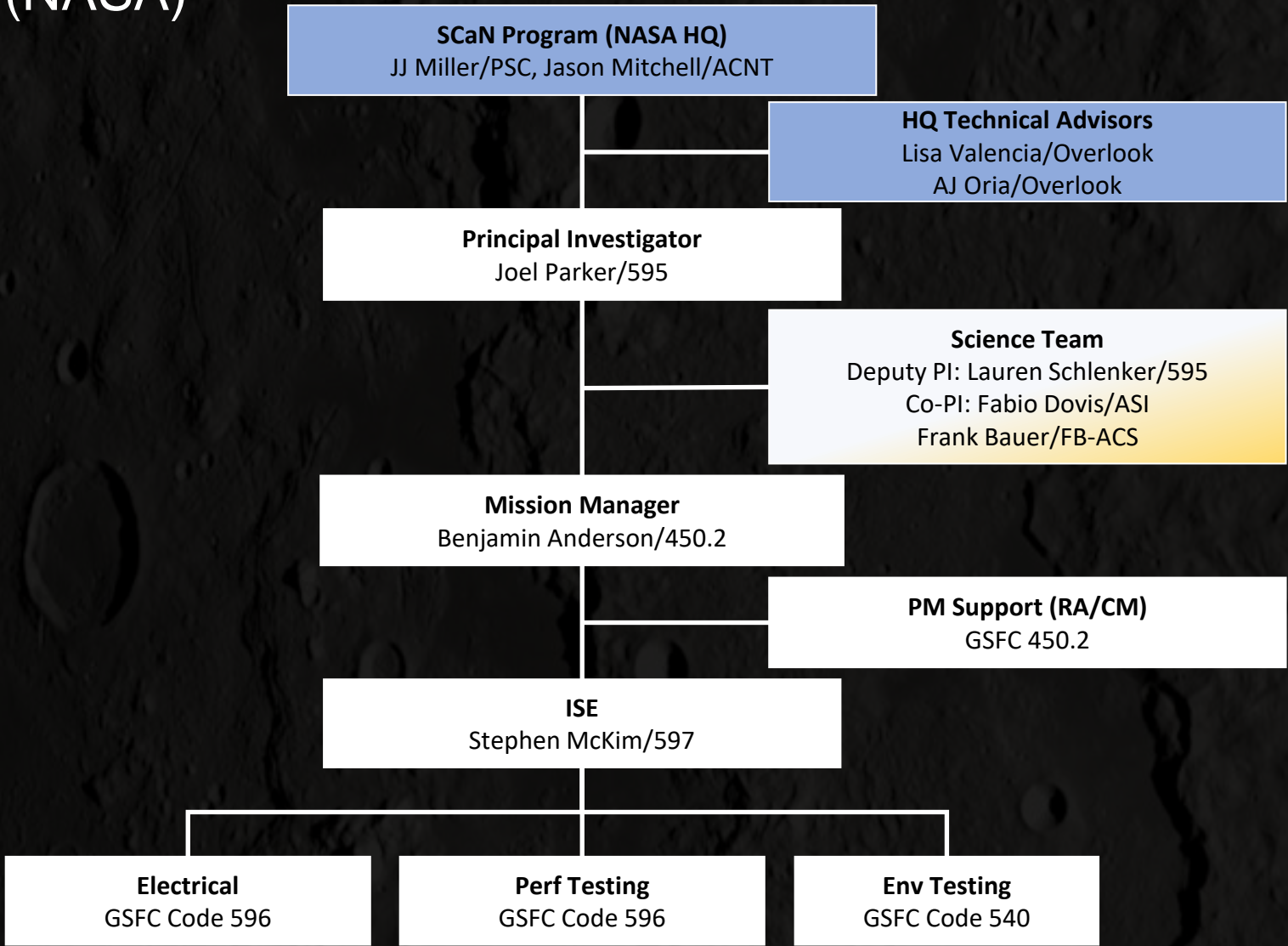
LunaNet



# LuGRE Project Organization (NASA)



NASA/HQ	NASA/GSFC
Firefly	ASI



# Integration & Test

